

Abstract

Abstract: As a solution for the lost-in-space star identification problem we present Star Identification using Graph Neural Network (SIGNN), a novel approach using Graph Attention Networks. By representing the celestial sphere as a graph data structure, created from the ESA's Hipparcos catalogue, we are able to accurately capture the rich information and relationships within local star fields. Graph learning techniques allow our model to aggregate information and learn the relative importance of the nodes and structure within each stars local neighbourhood to it's identification. This approach, combined with our parametric data-generation and noise simulation, allows us to train a highly robust model capable of accurate star identification even under intensive noise, outperforming existing methods.

Introduction

Motivation:

- Celestial navigation as an offline alternative to GPS, which is vulnerable to interference (spoofing, jamming, outages etc).
- Star detection is largely solved, but star recognition remains challenging in noisy conditions
- Graph Neural Networks (GNNs) can capture complex relationships in graph data

Contribution:

- SIGNN uses graph-based representations of star-fields + noise simulations to learn robust star representations for identification.
- Achieves superior accuracy over classical and ML-based methods, even at high noise.

Existing Work & Celestial Sphere

Classical:

- Sub-graph isomorphism (Triangle, pyramid)
- Star pattern recognition (Grid, radial)

Machine Learning:

- CNN-based: spider-web images.
- Representation learning: RPNet w/ autoencoders

Celestial Sphere (RA/DEC):

- Stars mapped via Right Ascension (RA) and Declination (DEC) analogous to longitude/latitude on Earth
- Celestial coordinates used to compute angular separations (edges) among connected stars in a graph representation that preserves exact geometry and precision.



Methodology – Graph Construction

Celestial Sphere Graph Nodes & Edges:

- Star data extracted from Hipparcos catalogue
- Nodes = stars (RA, DEC, V_{mag}).
- Edges = angular distance (Vincenty) $< \tau$ (3°).

Filtering:

- Exclude stars fainter than magnitude 6.0 (visible to naked eye / simulated sensor threshold).
- Remove isolated nodes (fewer than 2 neighbours).

Source Detection:

- Accurate absolute V_{mag} is challenging; use relative brightness among direct neighbours (min-max).

Resulting Input Graph:

- Each node's feature = local relative magnitude (brightness) compared to direct neighbours).
- Edges represent angular distance, min-max scaled.
- Pattern radius = 3° ; 2-hop coverage = 6° .
- Final graph size: 4274 nodes, 20K edges.

SIGNN – Star Identification using Graph Neural Networks

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SIGNN Pipeline:

- 2 GAT layers \rightarrow linear classifier
- Multi-noise training: positional, dropped, false, magnitude
- Nodes aggregate from neighbours' features and edge features (angular distances)
- 2-hop receptive field of 6°



Figure: SIGNN architecture

Example: Sirius neighbourhood

Sirius (2-Hop neighbourhood):

Sirius is connected to its neighbours within 3°, then those neighbours connect further within 3°. The GAT layers aggregate these relationships to correctly identify Sirius despite noise. Below, we show two perspectives: (a) the immediate 2-hop neighbourhood, and (b) a 10° FOV subgraph centred on Sirius, showing an example sensor input.



(a) 2-Hop around Sirius, only connected nodes are (b) 10° FOV subgraph, simulated sensor input used in the representation of Sirius

Figure 1. Two visual perspectives of Sirius: a closer 2-hop neighbourhood vs. a wider 10° view.

Real World vs Noise Simulation

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 Positional: shift star RA/DEC by random angular distance offsets 	Sensor Issue	Detection Output Noise Simulation	
	Weather	Missing stars	Drop
 Dropped: randomly remove stars (cloud, occlusion) False: insert false sources (aircraft, sensor noise) Magnitude: perturb brightness by random % 		Wrong mag	Mag noise
	Occlusion	Missing stars	Drop
	Atmospheric Effects	Shift position	Position noise
		Vary mag	Mag noise
	Lens distortion	Shift position	Position noise
	Lights	Extra stars	False



In real images, detect centroids \rightarrow convert to angular measures.

Methodology – SIGNN Architecture

GAT (2-layer):

$$\mathbf{x}'_{i} = \prod_{h=1}^{H} \sigma \left(\sum_{j \in \mathcal{N}(i) \cup \{i\}} \alpha_{i,j}^{h} \mathbf{W}^{h} \mathbf{x}_{j} \right)$$

where $\alpha_{i,i}^h$ are learnable attention coefficients. Attention Mechanism:

- Each node attends to its neighbours
- Learnable weights $(\alpha_{i,i}^h)$ focus on important and reliable connections

Training Details:

- 500 epochs, batch size = 64
- Each epoch: 1024 variations of entire celestial sphere (20%) each noise + 20% no noise)
- Fast.
- No need for large, static stored dataset
- Bayesian optimization for hyperparams
- Each star predicted on \approx 480k times during training



Test Setup:

Comparisons: Classical (Grid), ML (RPNet, Spider).



Insights:

- SIGNN yields high accuracy for star in FOV under considerable noise conditions
- Graph representation and angular-distance approach is sensor-agnostic.

Noise Simulation & Training

Parametric dataloader generates noisy samples during training.

Figure 2. Four main noise types visualized.

Results & Noise Performance

• Generate random FOV-subgraphs (10°) across the celestial sphere, simulating sensor views with random boresight and location. • 20k images per noise type and level : 800k images, 28 million star predictions.

Figure 3. SIGNN outperforms prior methods under high noise.

Discussion & Future Work

Future:

- Edge-of-image problem: partial star patterns near FOV boundary.
- Extend to variable and fainter magnitude thresholds.
- Creation of real image dataset + testing

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- Start End Step Noise Type Position (Std Dev Radians) 0.0001 0.001 0.0001 1% 10% 1% Magnitude (%) 5% 50% 5% False (%) 5% 50% 5% Dropped (%)
- Table 2. Testing schema for noisy star recognition.